## ON A SOLUTION TO THE EQUATIONS OF MAGNETO-GASDYNAMICS

## (K RESHENIIU URAVNENII MAGNITNOI GAZODINAMIKI)

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An investigation of strong and weak discontinuities in magneto-hydro-dynamics is contained in a series of papers and books (see, for instance, [1-4]). In the following, the equations of planar flow in a magnetic field parallel to the velocity field are transformed under certain initial restraints to a linear equation of the Chaplygin type [5]. We will apply the result to a problem in which there are no strong discontinuities.

The equations of the steady motion of a gas with infinite conductivity in a magnetic field have the following form:

$$\operatorname{div} \mathbf{H} = 0, \quad \operatorname{rot} (\mathbf{W} \times \mathbf{H}) = 0, \quad \operatorname{div} \rho \mathbf{W} = 0, \quad (\mathbf{W} \cdot \nabla) \mathbf{W} = -\frac{\operatorname{grad} p}{\rho} - \frac{1}{4\pi\rho} \mathbf{H} \times \operatorname{rot} \mathbf{H}$$
 (1)

where H is the magnetic field strength, p,  $\rho$  and W are respectively the pressure, density and vector velocity of the flow. If the flow is planar and the vector H lies in the plane of the flow, it follows from the second of Equations (1) that W  $\times$  H = const. If W || H at one point, then W || H throughout the flow field. One can write

$$\mathbf{H} = k (x, y) \rho \mathbf{W} \tag{2}$$

where k(x, y) is the coefficient of proportionality.

From the first and third equations of the system (1) we conclude that k(x, y) = const along a streamline. The vector  $H \ddot{x}$  rot H is perpendicular to the streamline. Therefore, the Bernoulli formula

$$wdw + \frac{dp}{p} = 0 (3)$$

is correct along streamlines.

Let us assume  $p = p(\rho)$  and let formula (3) be correct in any direction in the region of flow. We will also consider subsequently that k = const throughout the flow, which obtains in particular for the undisturbed

parallel flow at infinity. On the basis of (3) we have

$$\rho = \rho(w), \qquad p = -\int \rho(w) w dw \tag{4}$$

from which there follows

$$\frac{1}{p} \operatorname{grad} p = -\operatorname{grad} \frac{w^2}{2} \tag{5}$$

On the other hand,  $(W \cdot \nabla) W = \text{rot } W \times W + \text{grad } \frac{1}{2} W^2$ . Therefore, the last equation of the system (1) reduces to the form

$$rot W \times W = -\frac{1}{4\pi\rho} \mathbf{H} \times rot \mathbf{H}$$
 (6)

Projecting Equation (6) on to the coordinate axes x, y and taking formula (2) into account, we obtain

$$\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} = \frac{k^2}{4\pi} \left( \frac{\partial \rho v}{\partial x} - \frac{\partial \rho u}{\partial y} \right), \quad \text{или} \quad \frac{\partial v^*}{\partial x} - \frac{\partial u^*}{\partial y} = 0$$
 (7)

where

$$u^* = w^* \cos \theta, \qquad v^* = w^* \sin \theta, \qquad w^* = w \left(1 - \frac{k^2}{4\pi} \rho\right)$$
 (8)

and  $\theta$  is the angle of the velocity vector with the abscissa.

The existence of a stream function  $\psi(x, y)$ 

$$\frac{\partial \psi}{\partial x} = -v\rho(w) = -v^{\bullet}\rho^{\bullet}(w^{\bullet}), \qquad \frac{\partial \psi}{\partial y} = u\rho(w) = u^{\bullet}\rho^{\bullet}(w^{\bullet}) \qquad \left(\rho^{\bullet} = \frac{\rho}{1 - k^{2}\rho/4\pi}\right)$$
(9)

follows from the continuity equation.

Equation (7) permits a fictitious potential  $\phi$  to be introduced in accordance with the formula

$$\frac{\partial \varphi}{\partial x} = u^*, \quad \frac{\partial \varphi}{\partial y} = v^* \tag{10}$$

As is well-known from the equations of total differential expressions

$$dx = \frac{\cos \theta}{w} d\phi - \frac{\sin \theta}{\rho w} d\psi, \qquad dy = \frac{\sin \theta}{w} d\phi + \frac{\cos \theta}{\rho w} d\psi \tag{11}$$

one can derive the following system of equations for the unknown function  $\phi$  and  $\psi$ :

$$\frac{\partial \varphi}{\partial \theta} = \frac{w^{\bullet}}{\rho^{\bullet}} \frac{\partial \psi}{\partial w^{\bullet}}, \quad \frac{\partial \varphi}{\partial w^{\bullet}} = w^{\bullet} \frac{d}{dw^{\bullet}} \left(\frac{1}{\rho^{\bullet} w^{\bullet}}\right) \frac{\partial \psi}{\partial \theta}$$
 (12)

The system (12) has the canonical form

$$\frac{\partial \varphi}{\partial \theta} = \sqrt{K} \frac{\partial \psi}{\partial s}, \qquad \frac{\partial \varphi}{\partial s} = -\sqrt{K} \frac{\partial \psi}{\partial \theta}$$
 (13)

where the functions of the velocity  $\sqrt{K}$  and s are related to  $w^*$  and  $\rho^*$  in

accordance with (6) by the formulas

$$\frac{|dQ|}{ds} = -V\overline{K}P, \qquad Q = V\overline{K}\frac{dP}{ds} \qquad \left(P = w^{*-1}, \ Q = (\rho^*w^*)^{-1}\right) \tag{14}$$

hence

$$ds = \left(\frac{P'\left(w^{\bullet}\right)Q'\left(w^{\bullet}\right)}{P\left(w^{\bullet}\right)Q\left(w^{\bullet}\right)}\right)^{1/2}dw^{\bullet}, \qquad V\overline{K} = \left(\frac{Q\left(w^{\bullet}\right)Q'\left(w^{\bullet}\right)}{P\left(w^{\bullet}\right)P'\left(w^{\bullet}\right)}\right)^{1/2}$$
(15)

Substituting expressions for the functions  $P(w^*)$  and  $Q(w^*)$  into (15) and taking into consideration formulas (8) and (9) and the formula for determining the velocity of sound

$$a^2 = \frac{dp}{d\rho} = -\frac{w\rho(w)}{\rho'(w)} \tag{16}$$

we obtain

$$V\overline{K} = \frac{1}{\rho} \left( \frac{(1-M^2)(1-m\rho)^8}{1-m\rho(1-M^2)} \right)^{1/2}, \quad ds = \pm \left( \frac{(1-M^2)[1-m\rho(1-M^2)]}{1-m\rho} \right)^{1/2} \frac{dw}{w} (m = k^2/4\pi)$$

where M is the Mach number. The negative sign in (17) is taken for the interval of variation of w in which  $dw^*/ds < 0$ . For imaginary values of s and  $\sqrt{K}$  we have the hyperbolic system of equations

$$\frac{\partial \varphi}{\partial \theta} = \sqrt{\chi} \frac{\partial \psi}{\partial \sigma} , \quad \frac{\partial \varphi}{\partial \sigma} = \sqrt{\chi} \frac{\partial \psi}{\partial \theta} \qquad (\sigma = -is, \sqrt{\chi} = -i \sqrt{K})$$
 (18)

In the case  $p = \text{const } \rho^{\kappa}$ , where  $\kappa$  is the ratio of specific heat coefficients, formulas (17) and (21) take the form

$$V\overline{K} = \left(\frac{(1-\lambda^2)\left[1-k_1\left(1-\lambda^2/h^2\right)^{\gamma}\right]^3}{(1-\lambda^2/h^2)^{h^2}\left[1-k_1\left(1-\lambda^2/h^2\right)^{\gamma}\left(1-M^2\right)\right]}\right)^{1/2}$$

$$ds = \pm \left(\frac{(1-\lambda^2)\left[1-k_1\left(1-\lambda^2/h^2\right)^{\gamma}\left(1-M^2\right)\right]}{(1-\lambda^2/h^2)\left[1-k_1\left(1-\lambda^2/h^2\right)^{\gamma}\right]}\right)^{1/2}\frac{d\lambda}{\lambda}$$

$$\left(k_1 = \frac{k^2}{4\pi}\left(\frac{\varkappa+1}{2\varkappa}a_{\bullet}^2\right)^{\gamma}, h^2 = \frac{\varkappa+1}{\varkappa-1}, \gamma = \frac{1}{\varkappa-1}, \varkappa\neq 1\right)$$

where  $\lambda$  is the magnitude of the relative velocity and  $a_*$  is the critical velocity of sound. If  $k_1 < 1$ , the system (13) is correct for  $\lambda < 1$  and the system (18) is correct for  $\lambda > 1$ . The case  $k_1 > 1$  is of greater interest. Then the quantities

$$1-k_1(1-M^2)(1-\lambda^2/h^2)^{\gamma}, \qquad 1-k_1(1-\lambda^2/h^2)^{\gamma}$$

which are negative in the neighborhood of  $\lambda=0$  vanish respectively for  $\lambda_1<1$  and  $\lambda_2=h(1-k_1^{1-\kappa})$ , where  $\lambda_2>\lambda_1$ . In the interval of velocity variation  $0<\lambda<\lambda_1$  we have for every  $\lambda_2$  the elliptic system of equations (13). If  $\lambda_2<1$ , we have for the subsonic interval  $\lambda_1<\lambda<\lambda_2$  the hyperbolic system of equations (18), and subsequently for the interval

 $\lambda_2 < \lambda < 1$  the elliptic system of equations (13), and finally for  $\lambda > 1$  again the system (18). If  $\lambda_2 > 1$ , the system (18) is correct for the intervals  $\lambda_1 < \lambda < 1$  and  $\lambda_2 < \lambda < h$ , and the elliptic system of equations (13) for the supersonic interval  $1 < \lambda < \lambda_2$ . If

$$\lambda_1 = 1$$
, или  $k_1 = \left(\frac{h}{h-1}\right)^{\Upsilon}$ 

the system (13) is correct for the whole interval  $\lambda_1 < \lambda < h$ .

Extracting the principal parts of the formulas (19) in the neighborhood of the singular points  $\lambda_1$  and  $\lambda_2 \neq 1$ , we find that in the first case  $\sqrt{K} \approx \text{const } s^{-1/3}$  and in the second case  $\sqrt{K} = \text{const } s$ , where s is to becomputed respectively for  $\lambda_1$  and  $\lambda_2$ .

For k=0 we have the usual equations of Chaplygin. For separate formulas  $p=p(\varphi)$  and values  $k_1$  a system of equations in Legendre functions can be found which are more convenient for solution than the system (13).

We will pass from the functions  $\phi$ ,  $\psi$  to the functions  $\Phi$ ,  $\Psi$  by means of the Legendre transformation

$$\Phi = x \frac{\partial \varphi}{\partial x} + y \frac{\partial \varphi}{\partial y} - \varphi, \qquad \Psi = x \frac{\partial \psi}{\partial x} + y \frac{\partial \psi}{\partial y} - \psi$$
 (20)

We have

$$x = \Phi_{u^{\bullet}} = -\Psi_t, \qquad y = \Phi_{v^{\bullet}} = \Psi_r \qquad \left(r = \frac{\cos \theta}{Q}, t = \frac{\sin \theta}{Q}\right)$$
 (21)

In the independent variables s,  $\theta$  the system (21) has the following form (see, for instance, [7])

$$\frac{\partial \Psi}{\partial \theta} = -\sqrt{K_1} \frac{\partial \Phi}{\partial s}, \quad \frac{\partial \Psi}{\partial s} = \sqrt{K_1} \frac{\partial \Phi}{\partial \theta} \qquad \left(\sqrt{K_1} = \sqrt{K} \left[\frac{P}{Q}\right]^2\right)$$

Approximate and exact methods of solution of the Chaplygin equations can be used for the solution of problems of a given flow of gas in a magnetic field. For the approximations one must use a closure condition. For instance, if in the order of the approximation some other function f(s) is taken instead of the rigorous dependence on  $\bigvee K(s)$ , after substituting the function f(s) in place of  $\bigvee K$  in Equation (14) we obtain an equation for determining the functions P(s) and Q(s). We obtain the dependence of  $\rho$  on w in the parametric form  $\rho = \rho(s)$ , w = w(s) by means of the formulas  $w^* = P^{-1}$ ,  $\rho^* = P/Q$  and the formulas (8) and (9).

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